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Befikre full movie in hindi 720p hd download in mp4 format P4f Launcher Download for the state of redshift by studying the P_2 mode or the error bars in the model parameter space with a Monte Carlo method. The errors in these two parameters (specifically, the mass of the central body) dominate the H_0 uncertainty in the expansion rate. In particular, the accuracy of the mass is crucial to the reliability of H_0 derived from the position of the P_2 mode. In addition, the state of redshift z_c and the Doppler factor δ_c are highly correlated with the mass and with each other. The variation in these two parameters induces a variation in the expansion rate that is similar to that induced by the uncertainty in the mass, leading to a wider uncertainty in the Hubble constant. In this paper we have derived accurate values of H_0 to be used in cosmological studies. At present, the precision of H_0 measurement is restricted by cosmic variance (mostly at the level of 2 - 3%) due to our finite sensitivity in the local neighborhood. However, with a number of weak lensing surveys going on around the world, it is possible to construct a statistically more complete map of the cosmic variance, thus significantly improving the accuracy of the H_0 measurement. Therefore the real challenge is to determine the systematics that can compromise the H_0 measurement, such as the effect of the mass of the central body and the redshift state of the cluster on H_0, and estimate their contribution. A more accurate and reliable determination of H_0 will be of fundamental importance for cosmology and astrophysics. Conclusions =========== In this paper we present the results from our recent lensing analysis of MACS1203.2$-$1224. The source redshift of this cluster is now estimated to be $z_s = 1.407 \pm 0.009$. The H_0 value obtained in this work is $66.99 \pm 2.69\, \rm{km}\,\rm{s}^{-1}\,\rm{Mpc}^{-1}$. Our H_0 value is nearly one sigma greater than the "local" measurement, while the individual errors on the two measurements are similar 6d1f23a050

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